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Lee, Yong Bok; Park, Chang Hyun

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STUDY FOR THE PERFORMANCE PARAMETERS OF A  
FLIR SYSTEM

Yong Bok Lee



# NAVAL POSTGRADUATE SCHOOL

## Monterey, California



# THESIS

STUDY OF THE PERFORMANCE PARAMETERS OF A  
FLIR SYSTEM

by

Yong Bok Lee

and

Chang Hyun Park

June 1981

Thesis Advisor:

E. C. Crittenden

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T199396



REPORT DOCUMENTATION PAGE		READ INSTRUCTIONS BEFORE COMPLETING FORM
1. REPORT NUMBER	2. GOVT ACCESSION NO.	3. RECIPIENT'S CATALOG NUMBER
4. TITLE (and Subtitle) Study of the Performance Parameters of a FLIR Systems		5. TYPE OF REPORT & PERIOD COVERED Master's Thesis June 1981
7. AUTHOR(s) Yong Bok Lee and Chang Hyun Park		6. PERFORMING ORG. REPORT NUMBER
9. PERFORMING ORGANIZATION NAME AND ADDRESS Naval Postgraduate School Monterey, California 93940		8. CONTRACT OR GRANT NUMBER(s)
11. CONTROLLING OFFICE NAME AND ADDRESS Naval Postgraduate School Monterey, California 93940		10. PROGRAM ELEMENT, PROJECT, TASK AREA & WORK UNIT NUMBERS
14. MONITORING AGENCY NAME & ADDRESS (if different from Controlling Office)		12. REPORT DATE June 1981
		13. NUMBER OF PAGES 39
		15. SECURITY CLASS. (of this report)  UNCLASSIFIED
		15a. DECLASSIFICATION/DOWNGRADING SCHEDULE
16. DISTRIBUTION STATEMENT (of this Report)  Approved for public release; distribution unlimited		
17. DISTRIBUTION STATEMENT (of the abstract entered in Block 20, if different from Report)		
18. SUPPLEMENTARY NOTES		
19. KEY WORDS (Continue on reverse side if necessary and identify by block number)  NPS FLIR, Thermal Imaging, IR Imaging Performance Parameters		
20. ABSTRACT (Continue on reverse side if necessary and identify by block number)  The performance parameters of Detectivity and Noise Equivalent Temperature Difference for a Forward Looking Infra-Red (FLIR) system were measured after optimization of the system. The achieved performance approached the theoretically evaluated limiting values.		



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Study for the Performance Parameters of a  
FLIR System

by

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Submitted in partial fulfillment of the  
requirements for the degree of

MASTER OF SCIENCE IN PHYSICS

from the

NAVAL POSTGRADUATE SCHOOL  
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## ABSTRACT

The performance parameters of Detectivity and Noise Equivalent Temperature Difference for a Forward Looking Infra-Red (FLIR) system were measured after optimization of the system. The achieved performance approached the theoretically evaluated limiting values.



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## ACKNOWLEDGEMENT

We wish to express our appreciation to Professor Milne for his assistance and we wish to acknowledge the assistance of Mr. Bob Moeller and Mr. Ken Smith with the mechanical structure and electronic equipment.

We would like to especially thank Professor Eugene Crittenden whose wisdom and guidance were invaluable in preparation of this thesis.

Finally, we would like to express our grateful thanks to all Professors of the Physics Department.



## I. INTRODUCTION

### A. BACKGROUND

Thermal imaging systems are of recent origin. Since the first real-time FLIR was possible in late 1960, the development of the thermal imaging system has been very rapid in its design and technology. It was the direct result of the rapid development of various types of detectors. The superiority in performance in all-weather conditions, and at night made its application wider than expected.

FLIR sensor technology is making significant strides that could earn indispensable roles for these real-time, 24 hour sensors in future aircraft, remotely piloted vehicles (RPV), missiles and air defense systems. Even the immediate prospects for wider deployment of FLIR sensors could hinge on the recent efforts to reduce the cost, size, weight and complexity of the real-time imaging devices. Some types are in the testing phase in ASW, TRAM (Target Recognition Attack Multisensor) and Attack aircraft.

To provide a basis for better understanding of FLIR systems, a basic type of FLIR system was assembled. The several performance parameters were measured, but these values were at first, in poor agreement with the theoretical values. The principal effort has been in bringing theoretical and experimental values into agreement.



## B. OBJECTIVE

This paper is a brief review of the basic principles of operation and the performance parameter characteristics. Emphasis is placed on the measurement of  $D^*$  (Specific Detectivity) and the direct evaluation of NETD (Noise Equivalent Temperature Different) in terms of  $D^*$ .

The main scope of this study is in matching the measured and theoretically calculated NETD values. For this purpose the whole system was re-examined and several new experimental devices have been designed and constructed.



## II. THEORY

### A. SYSTEM PRINCIPLES

FLIR, Forward Looking Infra-Red, is a thermal imaging system which produces an image, primarily by self emission of a target and by the target's emissivity difference in the Infra-Red wavelength region.

The thermal radiation from a target in the IR region, 8-14  $\mu\text{m}$ , has good atmospheric transmission and also corresponds to the peak thermal radiation from an object at ambient temperature. The radiation is collected and focused onto an IR detector by an optical system. The impinging radiation on the detector excites charge carriers across the bandgap of a semiconductor.

For quantum detectors like HgCdTe, one excess electron-hole pair is created for each photon absorbed. For a photoconductive detector as used here, this changes the conductivity of the material. Since the attached circuit provides a constant current through the material; the voltage drop across the detector varies inversely with the incident radiation intensity. Naturally there is a lower limit to the photon energy below which the detector will not respond. The lower limit is the direct energy bandgap  $E_g$  and the wavelength corresponding to this lower photon energy limit is called the cut-off wavelength.



The detector converts the optical signal intensity into the analog electrical signal. This is then amplified and processed for display on a video monitor. One can reproduce the image of a target which generates Infra-Red thermal radiation even in the absence of ambient light. The whole system is described simply by the block diagram shown in Figure 1.

## B. SYSTEM PARAMETERS

Several detector and system parameters are used to describe the system performance. These are R, D\*, NEP, NETD, MRTD and MDTD, which individually represent; Responsivity, Specific Detectivity, Noise Equivalent Power, Noise Equivalent Temperature Difference, Minimum Resolvable Temperature Difference, and Minimum Detectable Temperature Difference. Some of the important parameters among these usually have the definition and expression given below.

### 1. NEP

Noise Equivalent Power (NEP) is defined as the signal power which produces a Signal-to-Noise ratio of unity.

$$NEP = \frac{P_d}{(V_s/V_n)} \quad [\text{watt}] \quad (1)$$

where  $P_d$  [watt] : signal power incident on the detector sensitive area, same as  $H_d A_d$



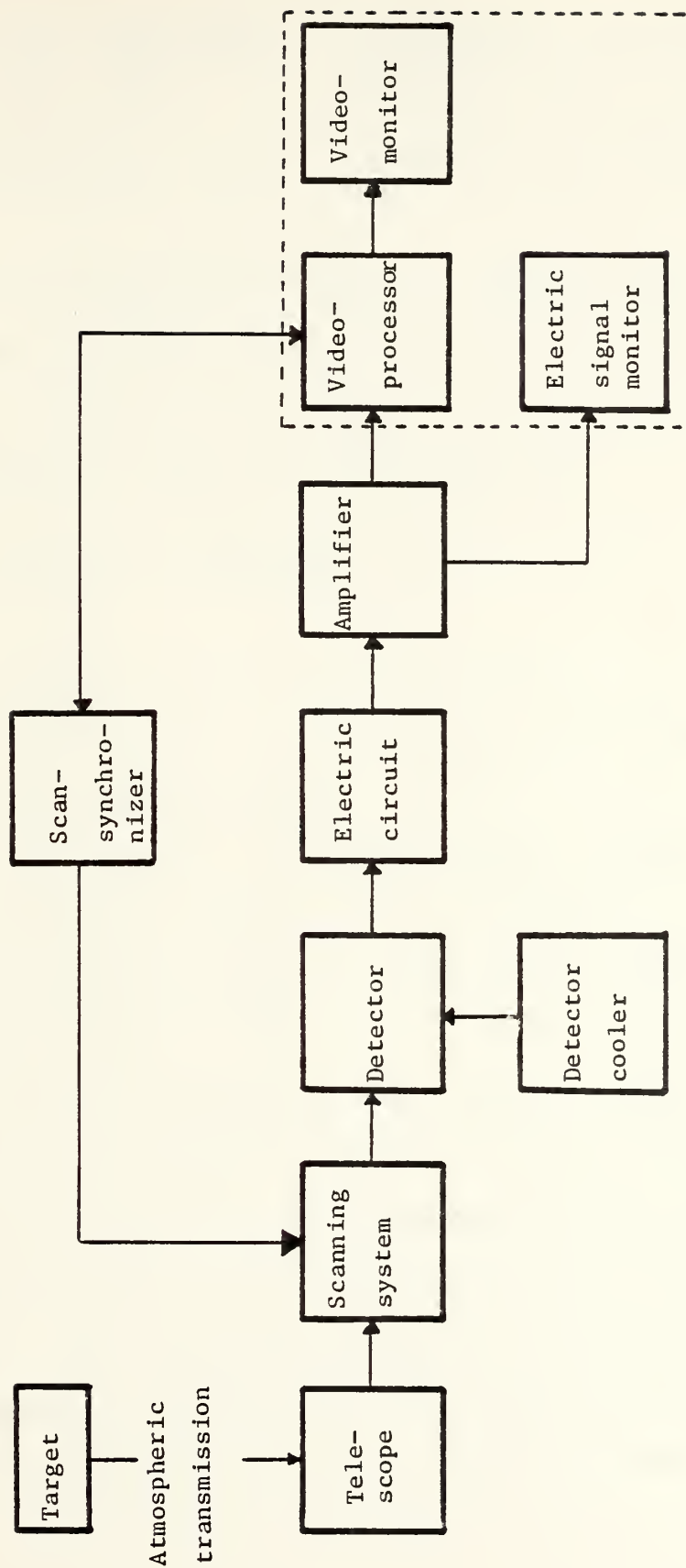


FIG. 1 FLIR system



$H_d$  [watt/cm<sup>2</sup>] : irradiance on the detector  
 $A_d$  [cm<sup>2</sup>] : detector sensitive area  
 $(V_s/V_n)$  : signal-to-noise voltage ratio

## 2. D\*

Specific Detectivity ( $D^*$ ) is a detector performance parameter which represents the detector output signal-to-noise ratio for one watt of input signal power for a unit detector area and a unit electrical bandwidth.

$$D^* (\lambda, f) = \frac{(A_d \Delta f)^{1/2}}{H_d A_d} (V_s/V_n) \quad (2)$$

where

$\Delta f$  [ Hz ] : system bandwidth

$V_n$  [volts] : detector rms noise voltage

measured in the bandwidth of  $\Delta f$ .

$V_s$  [volts] : detector signal voltage as a function of wavelength and electrical frequency.

$D^*$  comes from the expression of  $D$  (Detectivity) which is the reciprocal of NEP.

The assumptions made in the process of derivation require that detector noise varies as  $A_d^{1/2}$  and  $\Delta f^{1/2}$ , [Ref. 1]. The first assumption is valid for photon detectors as long as the detector areas do not vary by more than an order of magnitude. The second approximation requires either that noise is frequency invariant over the amplifier bandwidth, or that it is measured over such a narrow band that a



variation is insignificant. Further details can be found in Lloyd. [Ref. 2].

### 3. NETD

Noise Equivalent Temperature Difference (NETD) is a measure of the ability of a system to discriminate small signals in the presence of noise. NETD is defined as the blackbody target-to-background temperature difference in a standard test pattern which produces a peak signal-to-rms noise ratio of unity when the system views the test pattern. [Ref. 2]. This is given as

$$\text{NETD} = \frac{\Delta T}{(V_S/V_N)} \quad (3)$$

where

$\Delta T$  [ $^{\circ}\text{K}$ ] : temperature difference between a target and background

NETD is also theoretically derived by Lloyd [Ref. 2] as

$$\text{NETD} = \frac{\pi (A_d \Delta f)^{\frac{1}{2}}}{\alpha \beta A_c \int_0^{\infty} \frac{\partial W}{\partial t} D^*(\lambda) \tau_o(\lambda) d\lambda} \quad (4)$$

where

$\alpha, \beta$  : detector angular subtense which is  $\alpha = a/f, \beta = b/f$  where  $a, b$  are the size of a rectangular detector cell and  $f$  is the focal length of the optic system.

$\tau_o(\lambda)$  : IR optical transmission as a function of wavelength.

$W_\lambda$  [watt/cm<sup>2</sup>] : radiant emittance of a target



For some systems it is possible to assume that  $\tau_0(\lambda) = \text{const.}$  within specific wavelength region  $\lambda_1 \leq \lambda \leq \lambda_2$  and  $\tau_0(\lambda) = 0$  elsewhere.

Then

$$\text{NETD} = \frac{(A_d \Delta f)^{1/2}}{\alpha \beta A_o \tau_0} \frac{1}{\int_{\lambda_1}^{\lambda_2} \frac{\partial W_\lambda(T)}{\partial T} D^*(\lambda) d\lambda} \quad (5)$$

#### 4. MRTD

Minimum Resolvable Temperature Difference (MRTD) is defined as the image signal-to-noise ratio required for an observer to resolve an equally spaced four bar target that is obscured by noise. MRTD is also derived by Lloyd [Ref. 2] under several assumptions.

$$\text{MRTD} = \frac{3 (\text{NETD}/\Delta f) f_T (\alpha \beta)^{1/2} / \tau_d}{r_s (T_e \dot{F})^{1/2}} \quad (6)$$

where

$f_T$  [Hz/mrad] : fundamental target frequency  
 $T_e$  [sec] : effective eye integration time  
 $\tau_d$  [sec] : detector dwell time  
 $\dot{F}$  [Hz] : frame rate  
 $r_s$  : overall system MTF

#### 5. MDTD

Minimum Detectable Temperature Difference (MDTD) is the blackbody temperature difference required for an



observer to detect the presence of a square target when he is allowed unlimited time to make a decision and knows where to look for the target. Lloyd derived the expression of MDTD under the same assumption as MRTD.

$$\text{MDTD} = \frac{r_s 1.5 \sqrt{2} \text{MRTD} (f_T - 1/2w)}{\overline{I(x,y)}} \quad (7)$$

where

$\overline{I(x,y)}$  : average value of the convolution integral of the image of the square target

The derivations for the NETD, MRTD, and MDTD are carried out in detail with all assumptions in Lloyd [Ref. 2], so the derivations are omitted here.

The parameters, MRTD and MDTD, have important meaning in expressing the ability of the human eye to resolve or detect the target image on the video monitor. However, this is not a key point in this study.

D\* and NETD are by far the most frequently used figures of merit for characterizing the sensitivity of Infra-Red quantum detectors and FLIR systems. Since those two parameters are considered the most basic and fundamental parameters, further study of these was carried out.

### C. DIRECT EVALUATION OF NETD IN TERMS OF D\*

In order to compare measured and theoretical values of NETD, a form is needed which permits use of directly measured quantities. Calculation of the theoretical



value of NETD by the expression from Lloyd is a time consuming job. It also requires knowledge of the spectral variation of  $D^*$ , not available without additional equipment which is not presently available. A more convenient form of this parameter expression is developed as follows.

### 1. Derivation of $D^*$

Basically the expression for Specific Detectivity is developed for the ultimate derivation of NETD expression rather than for itself. To derive a new expression of  $D^*$  one can imagine a simple experiment set-up as shown in Figure 2.

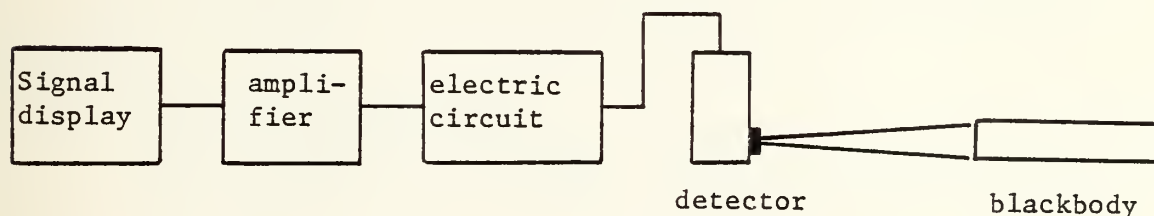


Fig. 2 Experiment set-up for detectivity derivation.

Total irradiance at the blackbody [watt]:  $H_S A_S$

where

$H_S$  [watt/cm<sup>2</sup>] : irradiance of the blackbody

$A_S$  [cm<sup>2</sup>] : area of blackbody perpendicular

to the line of sight



Total irradiance at the blackbody per steradian

$$[\text{watt}/S_r] : H_s A_s / \pi$$

Detector area expressed in terms of solid angle  $S_r$ :

$A_d/Z^2$  where  $Z$  [cm] : distance between detector and blackbody

Total power absorbed into detector  $P_d$

$$P_d [\text{watt}] = (A_d/Z^2) (H_s A_s / \pi) \quad (8)$$

General expression of  $D^*$

$$D^* [\text{cm } H_z^{1/2} / \text{watt}] = \frac{(A_d \Delta f)^{1/2}}{P_d} (V_s / V_n)$$
$$D^* = \frac{(A_d \Delta f)^{1/2} Z^2 \pi (V_s / V_n)}{A_d A_s H_s} \quad (9)$$

All terms in the expression for  $D^*$  in equation (9) can be directly read or measured in the experimental process.

## 2. Derivation of NETD

The expression for NETD, directly based on the definition of it, provides a good starting point to derive a more convenient form. This also leads to a form capable of direct experimental verification.

$$\text{NETD} = \frac{\Delta T}{(V_s / V_n)} \quad (10)$$

Figure 3 shows the simplified optical system, target and detector excluding all of the electronic elements. A pinhole (Area  $A_p$ ) is also attached to provide an ability to improve the spatial resolution.



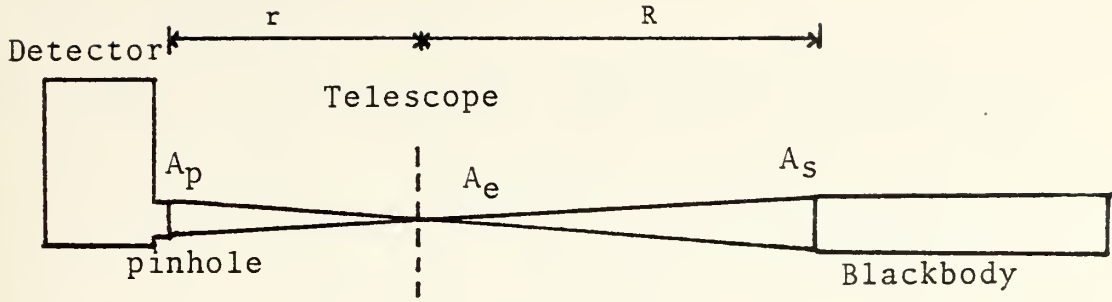


Fig. 3 Experiment set-up for NETD Derivation

Power absorbed by the detector  $P_d = A_e H_s A_s / R^2$  where  $R$  is the distance between the telescope primary mirror and a target

$$\text{Optics } f/\# = r/D = N$$

where  $D$  is a diameter of primary mirror

$$\text{Area of primary mirror: } A_e \text{ [cm}^2\text{]}$$

$$\text{Irradiance of blackbody surface: } H_s \text{ [watt/cm}^2\text{]} = G_s \Delta T$$

where  $G_s$  is a proportionality coefficient

Then,

$$P_d = (1/R^2) (1/4) \pi (r/N)^2 (A_s/\pi) G_t \Delta T \quad (11)$$

$$\text{where } \Delta T = 4N^2 P_d / A_p G_t$$

From Equation (2), (9), (10), (11)

$$\text{NETD} = \frac{4N^2}{A_p G_t} (A_d \Delta f)^{1/2} A_d A_s G_s \Delta T / (A_d \Delta f)^{1/2} Z^2 (V_s/V_n)$$

Finally we can get a much more convenient form of NETD, that does not involve specifying the wavelength limits.



$$\text{NETD} = \frac{4N^2 A_d A_s}{\pi A_p Z^2} (V_n/V_s) \Delta T \quad (12)$$

where  $G_s/G_t = 1$  if one keeps consistency in temperature. All terms appearing in this expression for NETD in Equation (12) can be directly read or measured experimentally. It is clear from the Equation (12) that if  $\Delta T$  is kept constant then NETD depends only on the pinhole size.



### III. EXPERIMENTAL PROCEDURE AND MEASUREMENT

#### A. APPARATUS USED FOR EVALUATION

The FLIR system to be evaluated is a HgCdTe single cell, mechanically scanning thermal imaging system assembled at NAVPGSCOL. Since a former study [Ref. 3] described precisely the whole system, only the main feature of elements are briefly discussed here.

##### 1. Optics; Cassegranian Type Dahl-Kirkham Telescope

- a. Diameter (D) : 15.24 cm
- b. Focal Length (f) : 228.6 cm
- c. Total Collecting Area : 172.8 cm<sup>2</sup>
- d. To see the effect of the pinhole on the value

NETD, an assembly of field lens and a pinhole are located in front of the detector.

e. Scanning mechanism which consists of 2 mirrors mutually perpendicular to each other is mounted at an angle of 45° to the beam exiting to the back end of the telescope. The scanning system could be located close to the exit aperture which means that it can be operating in the region where the image beam is parallel. The fact that a convergent beam scanner might cause the image deformation by mirror induced focal point shifting can be compensated. Because of the necessity for small mirror motion and the long focal length, the use of the convergent beam scanning technique



would not cause unacceptable image blur through beam focus.

## 2. Detector

a. Mercury Cadmium Telluride single cell detector with an IRTAN-2 window operates in the 8-14  $\mu\text{m}$  wavelength region.

b. A detector cell is mounted in a side-looking dewar and utilizes liquid nitrogen cooling down to 77  $^{\circ}\text{K}$  and is evacuated to  $2 \times 10^{-6}$  Torr.

c. The spectral response of the detector depends on the composition of the alloy and a wide range of performance is shown in Figure 4 from the supplier's specification [Ref. 4].

d. Manufacturer: Santa Barbara Research Center

e. Precise characteristics like Relative Responsivity vs. wavelength, Detectivity vs. Frequency, Responsivity vs. Frequency, Detectivity vs. Temperature can be referred to SBRC. [Ref. 4].

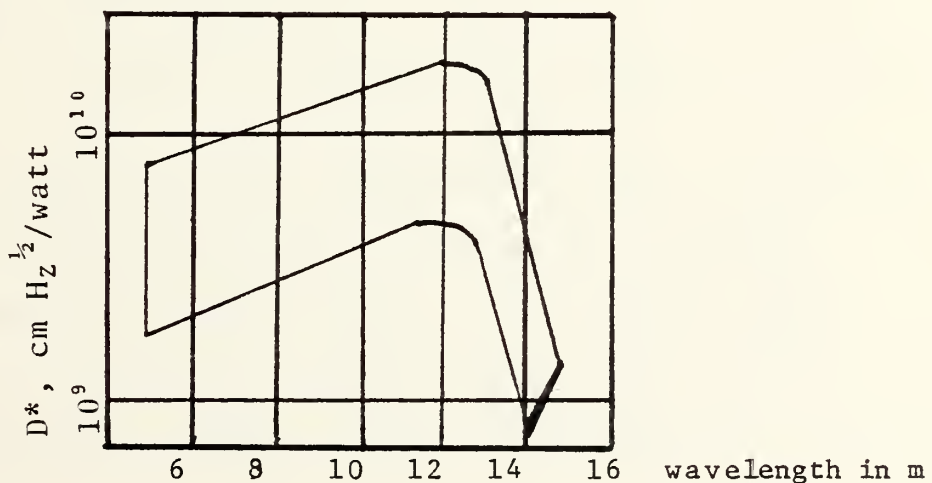


Fig. 4 Range of spectral detectivity



### 3. Blackbody Heat Source

A blackbody heat source is required for measuring the  $D^*$  and NETD. Initially a U-shaped heating element with a long aluminum bar was used as a blackbody with a powerstat to control the temperature of the heating element. But it was not good enough as a blackbody target due to the low emissivity of the target and very fast heat dissipation and severe fluctuation of temperature due to the ambient background.

So, a new blackbody target covered with asbestos and fiber glass as an insulator to prevent heat loss into the ambient background, was designed and built as shown in Figure 5.

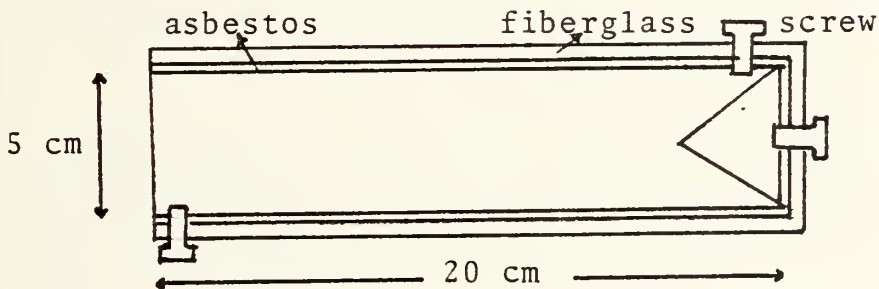


Fig. 5 Blackbody Heat Source

The cone inside this aluminum cylinder is used to increase internal reflections. Three screws at three different places are used to connect the thermocouples. A thermocouple attached to this blackbody heat source is used to measure the temperature with an accuracy of  $\pm 0.1$  °C.



#### 4. Electronic Equipment

The electronic equipment is described in detail in the former study. [Ref. 5] . A brief summary of the equipment follows:

- a. Hewlett Packard 3310 A Function Generator  
: Raster control
- b. General Scanning CCX 101 scanner control  
: mirror drive
- c. Wavetek 180 Function Generator  
: Raster control
- d. Princeton Applied Research Model 113 Preamplifier  
: detector signal amplifier
- e. Hewlett Packard 3400A RMS Voltmeter  
: noise voltage measurement
- f. Monsanto AM-6419/USM-368 OSC.  
: display

Since HgCdTe is used as a photoconductor, a special electric circuit is needed to provide the conduction current. Figure 6 shows this electric circuit.

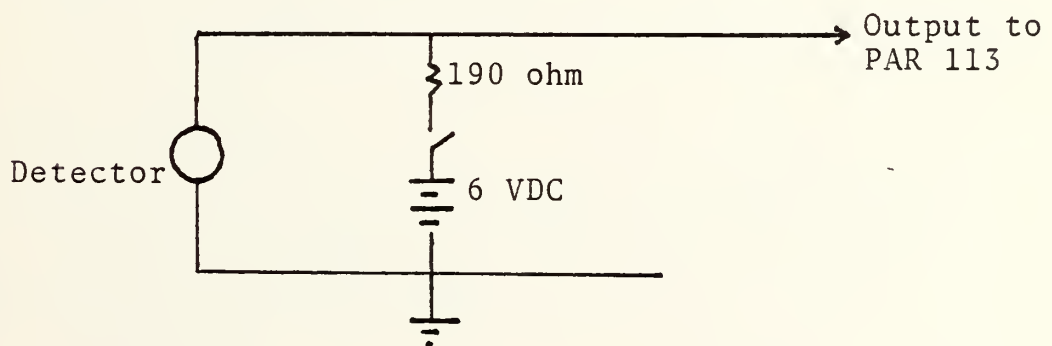


Fig. 6 Electric Circuit



## B. MEASUREMENT AND CALCULATION

### 1. D\*

To determine the detectivity it is necessary to measure the signal-to-noise voltage ratio. For this purpose a simple experiment is set up as shown in Figure 7.

Since the detector can detect only a change of radiation, a chopper is used to convert the steady radiation from the blackbody heat source into intermittent radiation. An aluminum plate, which has a hole in it, and is insulated with asbestos and fiber glass on both sides to prevent any thermal effect from the chopper or from any other heat source except blackbody heat source itself, is used as a real blackbody target. The fact that this aluminum plate with a hole, not the blackbody heat source, acts as a real blackbody can be easily verified by observing that the signal voltage is dependent on the distance between the aluminum plate and the detector, not on the distance from the detector to the blackbody heat source.

$$\Delta f = 3000 \text{ Hz}$$

$$A_d = 0.04 \text{ cm}^2$$

$$A_s = \pi \text{ cm}^2$$

$$Z = 50 \text{ cm}$$

$$H_s = 0.059 \text{ watt/cm}^2 \text{ using calculator program}$$

[Appendix] where the inputs are

$$\lambda_1 = 8 \text{ } \mu\text{m}$$

$$\lambda_2 = 14 \text{ } \mu\text{m}$$



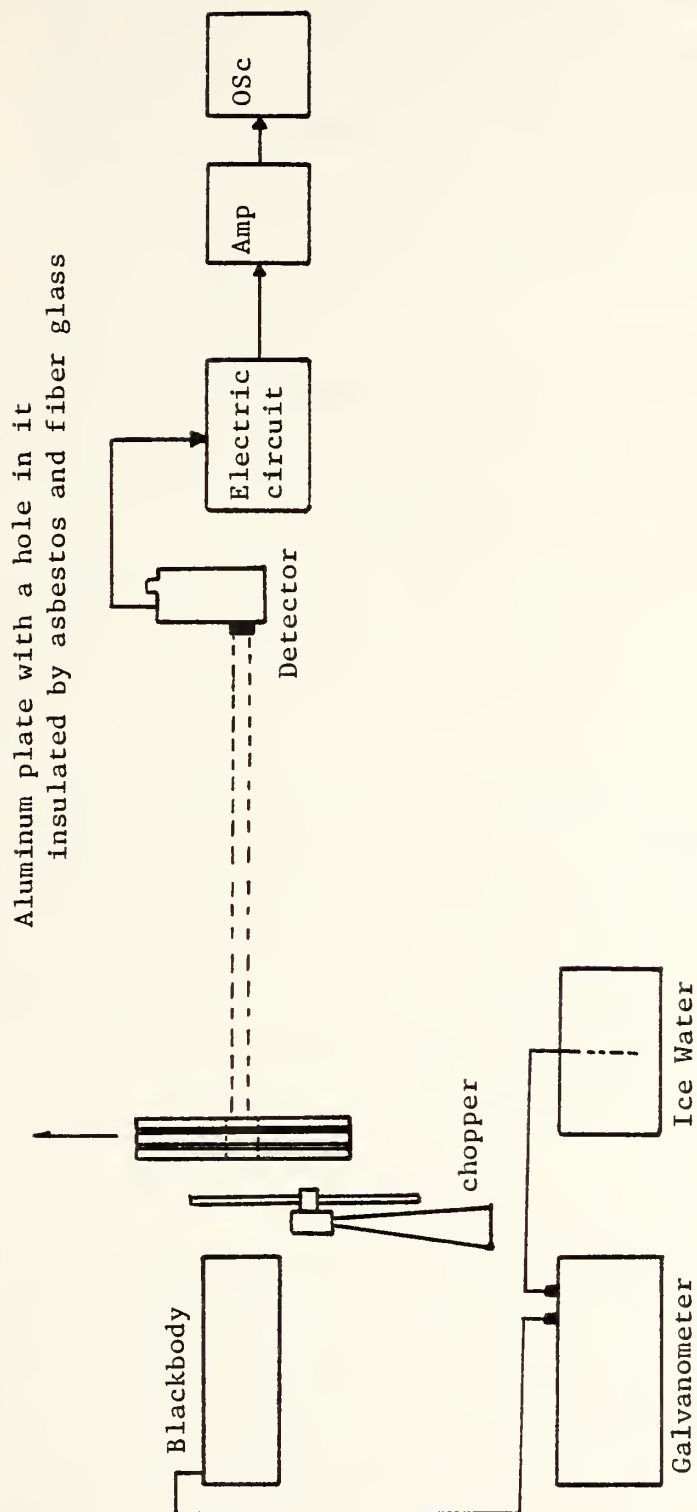


Fig. 7 Detectivity measurement experiment set-up



$T_t = 163\text{ }^{\circ}\text{C}$  (Blackbody Temperature)

$T_b = 22\text{ }^{\circ}\text{C}$  (Background Temperature)

$V_S/V_n = 400$  where measured value of

$V_S = 0.6\text{ volts}$  (Reading from the Oscilloscope)

$V_n = 0.0015$  (System Noise Voltage, reading from the rms Noise voltage meter)

Calculation of  $D^*$  using the previously derived detectivity expression, Equation (9) yields

$$D^* = 0.46 \times 10^{10} [\text{cmHz}^{\frac{1}{2}}/\text{watt}]$$

This value is the lower limit of detectivity in Figure 5 where the values are from the SBRC. [Ref. 4].

## 2. NETD

The experimental NETD can be measured, using the FLIR system under the same condition of blackbody temperature as for the  $D^*$  measurement by measuring the signal-to-noise voltage ratio for each size of pinhole. In this step it is important to ensure that the signal on the oscilloscope comes from the target, and not from any reflection.

The first thing to be checked was the shape of the image formed at the back focal point. An incandescent light bulb was used as a target at a long distance so that the incident light is considered as a parallel beam. Then, the image should be round and clear.

Secondly, the shape of a signal on the oscilloscope should have a flat top while turning the telescope



horizontally. This is based on the linear filter theory which map object distributions into image distributions by the process of convolution. This theory says that the image function is a weighted sum of the system response to the component delta functions of the object.

Every part of the system which might cause an image degradation due to the poor alignment and the reflection by system itself was checked. This procedure included checking the relative angle of the reflecting mirror to the beam exit from the telescope, the exact location of the pinhole at the back focal point, and the adjustment of the field lens assembly. Also an effort was made to keep the direction of the telescope scanning across the objective screen and the axis of the field lens-detector assembly on a horizontal path line.

Difficulty in keeping the target temperature constant was solved by using a newly designed blackbody heat source shown in Figure 5.

The chromel-alumel thermocouple was used with its reference junction in ice water.

Table 1 shows the measured signal and noise voltages and the calculated NETD based on this measurement using the Equation (3).



D of pinhole (cm)	$V_s$ volts	$V_n$ volts	$V_s/V_n$	NETD
0.0368	0.05	0.0015	33.3	4.23
0.05	0.09	"	60.0	2.35
0.75	0.135	"	90.0	1.57
0.1	0.188	"	125.3	1.12
0.15	0.241	"	160.7	0.87
no pinhole	0.38	0.0021	180.9	0.78

Table 1 NETD - Experimental Value



#### IV. ANALYSIS

Table 2 shows the theoretically calculated NETD and also the experimentally measured value as a function of pinhole size.

D. of pinhole	NETD		difference
	theoretical	experimental	
0.0368	3.77	4.23	0.46
0.05	2.04	2.35	0.31
0.075	0.91	1.67	0.76
0.1	0.51	1.12	0.61
0.15	0.23	0.87	0.64
no pinhole	0.1	0.78	0.68

Table 2 NETD - Theoretical and Experimental Value

The data of Table 2 are plotted in Figure 8. For reference, those values of NETD, measured and calculated in the former study [Ref. 5] are also included in Figure 8.

An analysis of the values of the performance parameter, NETD, in the table above, reveals several important points.

The experimentally measured values of the NETD are in reasonable agreement with the theoretical values calculated with Equation (11). Even though the difference is small enough, this discrepancy is understood as that, first, the temperature difference  $\Delta T$  still fluctuates on a certain scale which affects the measured signal-to-noise ratio.



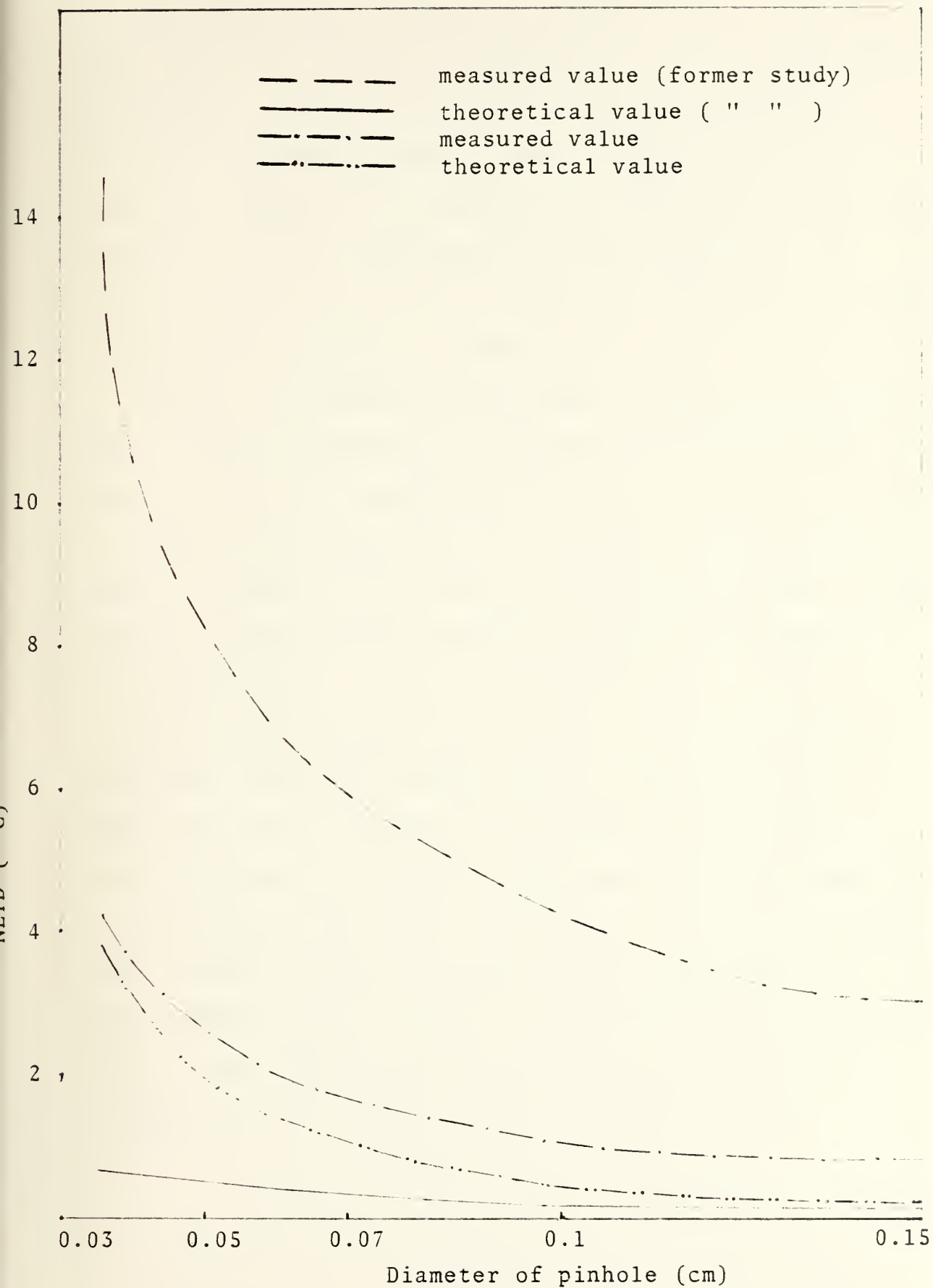


Fig. 8 NETD Comparison - theoretical and measured value



The temperature difference between the front end and back end of the blackbody furnace was 15 °C at the target temperature of 163 °C and ambient background temperature 22 °C. Considering that the approximate rate of change NETD with respect to the temperature difference ( $d \text{ NETD} / dT$ ) is about 0.02, the nonuniformity of the blackbody heat source can cause error in the NETD of 0.3 °C, which is half of the NETD difference in Table 2.

Second, this experiment basically assumed background limited noise conditions. But there should be another kind of noise, for example,  $1/f$  noise.

Third, the alignment problem is another factor decreasing the accuracy of the measured experimental value. Using the more precise way of aligning, the result is believed to be much better.

The expected dependency of the thermal sensitivity of the system expressed by NETD upon the various size of pinhole was well observed. As expected, as the pinhole size gets larger, the system shows more sensitive response to the thermal radiation. The smaller size of the pinhole shows better resolution. Consequently, when all of the other factors like  $T$ ,  $f/\#$  are fixed, the NETD is a function only of area of the pinhole. This relation is well shown in Figure 8.



## V. CONCLUSIONS

One of the basic performance parameters of the FLIR system, the NETD, was measured. Every effort was made to match the theoretical values and the experimentally measured values. For that purpose, a more convenient and direct form of NETD expression was derived. A new black-body heat source was designed to more closely approximate a true blackbody. Most of the laboratory effort was spent in establishing precise optical alignment and in keeping the target and the background temperature difference constant. As a consequence the discrepancy between the theoretical and experimental values of NETD decreased to 0.3-0.7 °C, close enough to be considered in agreement. The result of measuring and calculating the NETD for 5 different sizes of pinholes show the variation of NETD with the pinhole size to be as theoretically anticipated.

As a secondary objective, the detectivity was measured and found to be  $0.46 \times 10^{10} \text{ cm Hz}^{\frac{1}{2}}/\text{watt}$ . This value is the lower limit of detectivity in the characteristics provided by the detector manufacturer. Even though the value of  $D^*$  is somewhat lower than expected, it can be understood to be a quality decrease due to a long term use, and one can say that  $D^*$  is still in a tolerable region.



# APPENDIX

## CALCULATING "IN BAND" FLUX DENSITY OF BLACKBODY SOURCES USING TI-59

To use: press RST

enter  $\lambda_1$  microns R/S

enter  $\lambda_2$  " R/S

enter T A for °K

B for °F

C for °C

enter number of intervals

D for watt/cm<sup>2</sup>

E for photons/sec/cm<sup>2</sup>

000	42	STO	021	00	00	041	69	OP
001	01	01	022	69	OP	042	06	06
002	99	PRT	023	06	06	043	69	OP
003	91	R/S	024	69	OP	044	00	00
004	42	STO	025	00	00	045	75	-
005	02	02	026	91	R/S	046	03	3
006	99	PRT	027	76	LBL	047	02	2
007	91	R/S	028	12	B	048	95	=
008	76	LBL	029	42	STO	049	65	x
009	11	A	030	00	00	050	05	5
010	42	STO	031	02	2	051	55	÷
011	00	00	032	01	1	052	09	9
012	02	2	033	00	0	053	85	+
013	06	6	034	00	0	054	02	2
014	00	0	035	00	0	055	07	7
015	00	0	036	00	0	056	03	3
016	00	0	037	69	OP	057	95	=
017	00	0	038	04	04	058	42	STO
018	69	OP	039	43	RCL	059	00	00
019	04	04	040	00	00	060	91	R/S
020	43	RCL						



061	76	LBL	111	08	8	161	01	1
062	13	C	112	93	.	162	00	0
063	42	STO	113	03	3	163	07	7
064	00	00	114	02	2	164	04	4
065	01	1	115	03	3	165	52	EE
066	05	5	116	03	3	166	04	4
067	00	0	117	04	4	167	42	STO
068	00	0	118	55	÷	168	08	08
069	00	0	119	43	RCL	169	95	5
070	00	0	120	07	7	170	42	STO
071	69	op	121	55	÷	171	09	09
072	04	04	122	43	RCL	172	01	1
073	43	RCL	123	00	00	173	22	INV
074	00	00	124	54	)	174	23	LNx
075	69	OP	125	54	)	175	42	STO
076	06	06	126	75	-	176	06	06
077	69	OP	127	01	1	177	36	PGM
078	00	00	128	54	)	178	09	09
079	43	RCL	129	54	)	179	14	D
080	00	00	130	35	1/X	180	22	INV
081	85	+	131	65	x	181	52	EE
082	02	2	132	43	RCL	182	99	PRT
083	07	7	133	08	08	183	98	ADV
084	03	3	134	54	)	184	91	R/S
085	95	=	135	92	RTN	185	76	LBL
086	42	STO	136	76	LBL	186	15	3
087	00	00	137	14	D	187	98	ADV
088	91	R/S	138	98	ADV	188	42	STO
089	76	LBL	139	42	STO	189	05	05
090	16	A'	140	05	05	190	99	PRT
091	42	STO	141	99	PRT	191	98	ADV
092	07	07	142	98	ADV	192	53	(
093	53	(	143	53	(	193	43	RCL
094	53	(	144	43	RCL	194	02	02
095	43	RCL	145	02	02	195	75	-
096	07	07	146	75	-	196	43	RCL
097	45	Y <sup>x</sup>	147	43	rc1	197	01	01
098	43	RCL	148	01	01	198	54	)
099	09	09	149	54	)	199	55	÷
100	65	x	150	55	÷	200	43	RCL
101	53	(	151	43	RCL	201	05	05
102	53	(	152	05	05	202	95	=
103	43	RCL	153	95	=	203	42	STO
104	06	06	154	42	STO	204	03	03
105	45	Y <sup>x</sup>	155	03	03	205	01	1
106	53	(	156	03	3	206	93	.
107	01	1	157	93	.	207	08	8
108	04	4	158	07	7	208	08	8
109	03	3	159	04	4	209	03	3
110	08	8	160	01	1	210	06	6



211	05	5
212	01	1
213	06	6
214	52	EE
215	02	2
216	03	3
217	42	STO
218	08	08
219	04	4
220	42	STO
221	09	09
222	01	1
223	22	INV
224	23	LNx
225	42	STO
226	06	06
227	36	PGM
228	09	09
229	14	D
230	22	INV
231	52	EE
232	99	PRT
233	98	ADV
234	91	R/S



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